

## Closing: Controversies in UHP Tectonism

- maximum pressures
- maximum temperatures
- rates of exhumation
- sizes of UHP terranes

## Erosion Rates

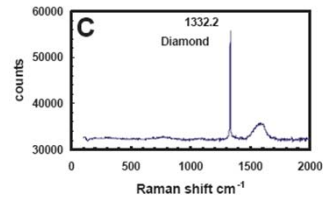
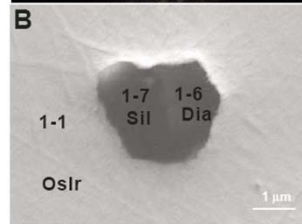
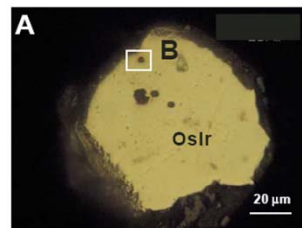
TABLE 2. Erosion rates by surface processes.

Burbank (2002)

Erosion process	Location	Measured	Rate of	Reference
Bare bedrock weathering	Wind River Range, Wyo.	Cosmogenic isotope abundance	5–20 m/m.y.	Small <i>et al.</i> (1997) Bierman (1994)
Rock-to-regolith conversion	San Gabriel Mtns, Calif.	Cosmogenic isotope abundance	≈350 m/m.y.	Heimsath (1999)
River incision	Indus River, Pakistan	Cosmogenic isotope abundance	5–10 km/m.y.	Burbank <i>et al.</i> (1996)
River incision	Himalayan Foreland	Deformed, dated terraces	10 mm/y	Lavé and Avouac (2000)
Bedrock landsliding	Southern Alps, New Zealand	Frequency magnitude data	5–15 km/m.y.	Hovius <i>et al.</i> (1997)
Glacial erosion	Alaska, New Zealand, Asia	Sediment volumes, reported rates	1–30 km/m.y.	Hallet <i>et al.</i> (1996)
Glacial erosion	Nanga Parbat	Glacial sediment load	5–7 mm/y	Gardner and Jones (1993)

## Luobusa, Tibet

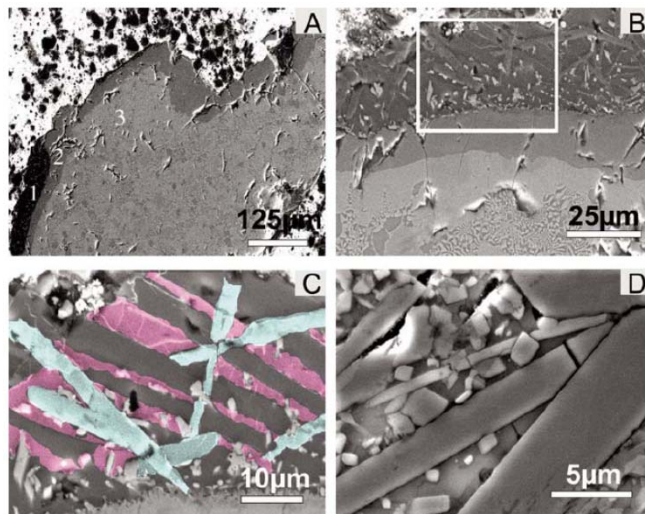
- diamonds in Oslr alloy
- coesite rimming FeTi alloy
- from chromitite



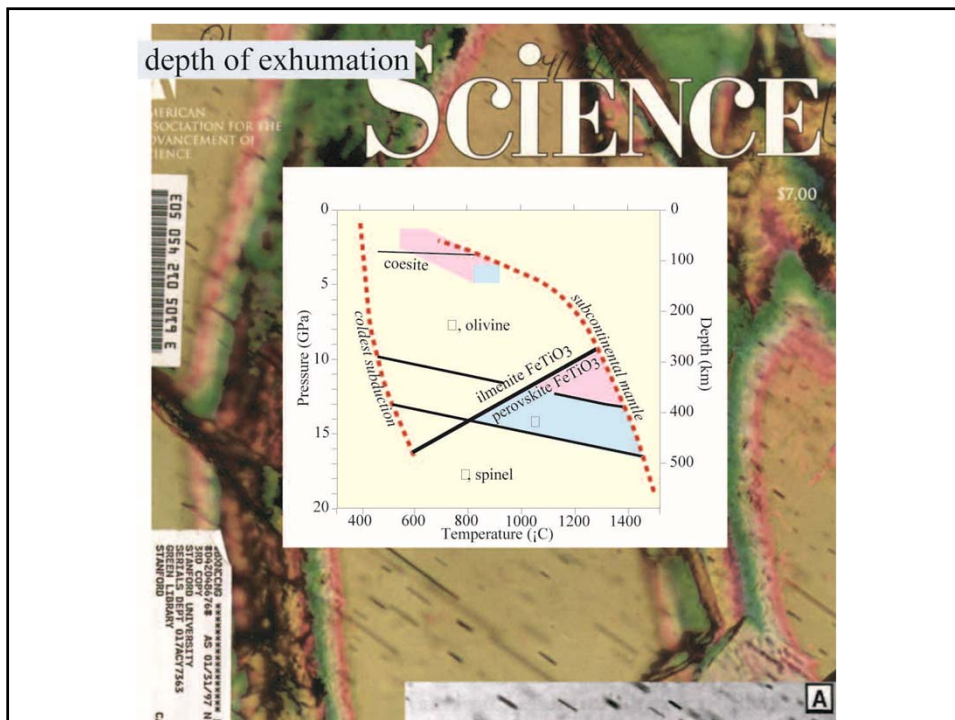
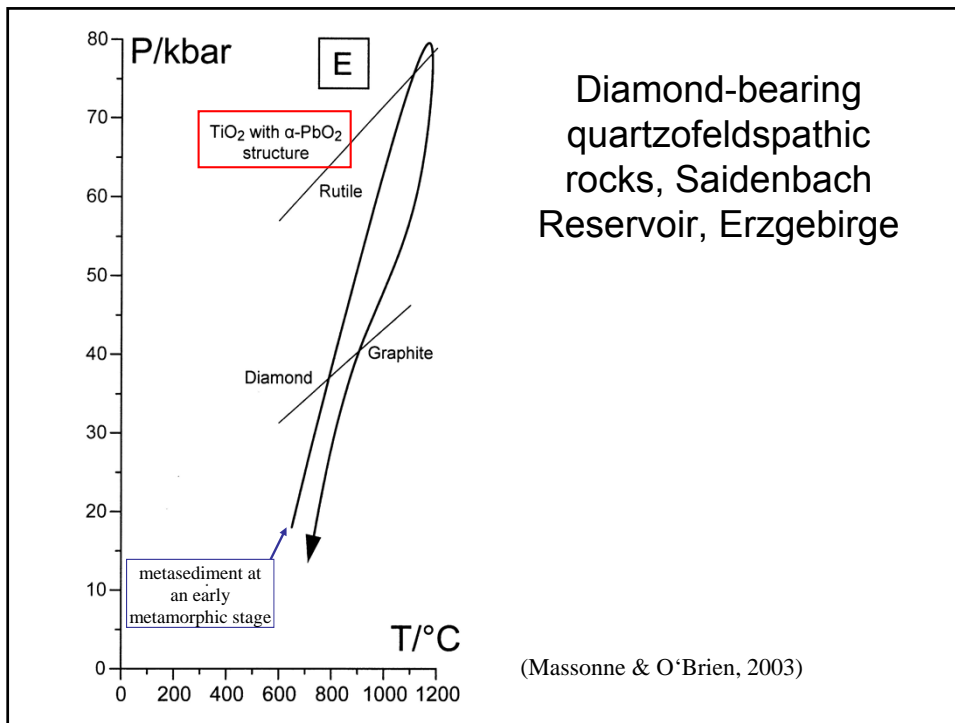
Yang et al. (2007)

## Luobusa

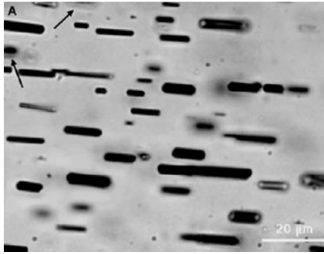
kyanite contains <12 wt% TiO<sub>2</sub>  
coesite crystals may be stishovite pseudomorphs



Yang et al. (2007)

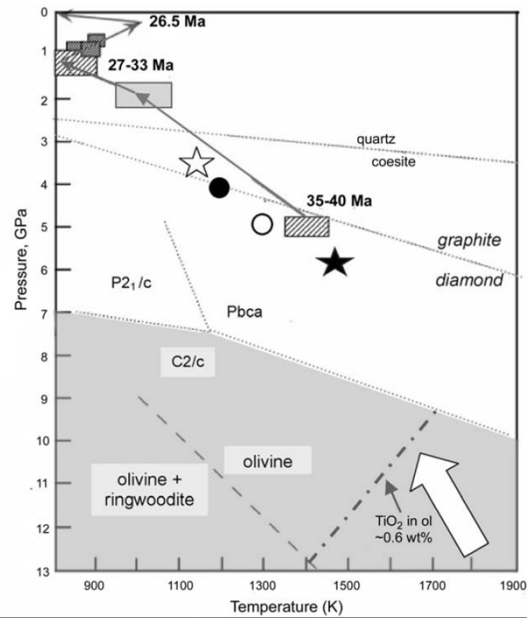


## Alpe Arami, Ti rods



defected olivine →  
ilmenite + chromite

Green (2005)



## Ti rods

1–3 vol%  $\text{FeTiO}_3$  rods in olivine, or 0.7–2.0 wt%  $\text{TiO}_2$  (Dobrzhinetskaya *et al.*, 1996) revised to 0.2–0.9 wt%  $\text{TiO}_2$  (Bozhilov *et al.*, 2005).

Ilmenite rods identical in size, composition & quantity to Alpe Arami, detected in Sulu (Hacker *et al.*, 1997) & Adula–Cima Lunga nappe (Risold *et al.*, 1999)

highest  $\text{TiO}_2$  content in kimberlite olivine < 0.06 wt% (Hervig *et al.*, 1986)  
0.11 wt%  $\text{TiO}_2$  in olivine at 10 GPa/1400 °C (Ulmer *et al.*, 1998)

$\text{TiO}_2$  = 0.035 wt% from EPMA measurements (Hacker *et al.*, 1997)

$\text{TiO}_2$  = 0.035 wt% from LA-ICP-MS (Risold *et al.*, 1996 & Reusser *et al.*, 1998)

## Ti rods

3 new polymorphs of  $\text{FeTiO}_3$  interpreted as intermediate between HP perovskite & ilmenite. Inferred exsolution during transformation of wadsleyite, which has high Ti solubility (Dobrzhinetskaya *et al.*, 1996)

diffraction patterns can be explained by dynamical diffraction between ilmenite & olivine matrix (Hacker *et al.*, 1997)

TEM shows rods controlled by presence of humite-type layers, *i.e.*,  $(\text{Mg}_2\text{SiO}_4)\cdot\text{Mg}(\text{OH},\text{F})_2$ . In such layers up to 50 mol%  $\text{MgH}_2$  replacement by  $\text{Ti}^{4+}$  is possible. (Risold *et al.*, 1999)